## The Vertical Structure of the Jovian Ring

J. A. Burns, L. Schaffer (Cornell University), D. P. Hamilton (U. Maryland), M. R. Showalter (Stanford University)

The ongoing Galileo imaging of the Jovian system will provide an opportunity to clarify the ring's radial and vertical structure, as well as the size distribution of its constituents, each of which is poorly constrained by Voyager images. Most models of the Jovian ring consider that electromagnetic forces elevate micron-sized grains (which acquire charge from the ambient plasma) out of the equatorial plane. These forces may explain one aspect of the ring system that stands out clearly: the pronounced vertical flaring of the ring in the halo region.

The current state of the observations is not sufficient to distinguish between competing models, and thus we review the various mechanisms proposed to date, highlighting their similarities and differences.

Here we use numerical orbit integrations over a range of particles sizes to follow dust grains originating from the main ring as they evolve by plasma drag through so-called Lorentz resonances. For a range of charge-to-mass ratios and plasma drag rates, large jumps in inclination ensue; these jumps occur near 1.7 and 1.4  $R_{\rm J}$ , where the vertical thickness of the main ring increases by an order of magnitude to the halo, and where the halo vanishes. This model does not depend strongly on plasma measurements in the ring region: all that is required is some small charge on grains, and any mechanism which gradually draws grains through the resonance zones. From these integrations we produce ring cross-sections which roughly match the halo slice data of Showalter et al.

As part of this study, we have simulated the trajectories of individual particles having widely different charge-to-mass ratios. Not surprisingly, very highly charged grains are closely tied to the magnetic field, while large particles follow nearly Keplerian orbits. Intermediately charged grains have significant excursions, some leaving the system altogether; this occurs especially near Lorentz resonances. These characteristics seem to provide the correct clay from which to mold the jovian ring's morphology.

Abstract submitted for 1996 DPS meeting

Date submitted: LPI electronic form version 5/96

## Division for Planetary Sciences Abstract Form

DPS Category 17 Running #7483 Session 0.00
Invited Poster presentation X Title only
Have you received your Ph.D. since the last DPS meeting?  Yes No X
Is your abstract newsworthy, and if so, would you be willing to prepare a new release and be available for interviews with reporters?  Yes No Maybe X
Paper presented by Joseph A. Burns  Center for Radiophysics and Space Research Space Sciences Building Cornell University Ithaca NY 14853 USA Phone: (607) 255 7186 Fax: (607) 255 6354 Email: jab16@cornell.edu
Special instructions: Tue Aug 27 16:53:11 CDT 1996
Membership Status (First Author):
DPS-AAS Member X Non-Member
Student Member Student Non-Member
Is this your first DPS presentation? Yes No X
Sponsor: